

Prestack migration-driven 5D interpolation - Program **interpolation5d**



Computation flow chart

Theory

Inadequate sampling has always been a problem for seismic acquisition. Obstacles give rise to gaps during land seismic acquisition which can cause migration artifacts. Missing azimuths hinder AVAz analysis and prestack inversion, while missing offsets hinder AVO and prestack inversion analysis. To address these data sampling limitations, various methods of 5D interpolation have been applied to predict missing data in otherwise sparse seismic surveys. Ideally the result is a more uniform distribution, increasing the signal-to-noise ratio and suppressing acquisition footprint.

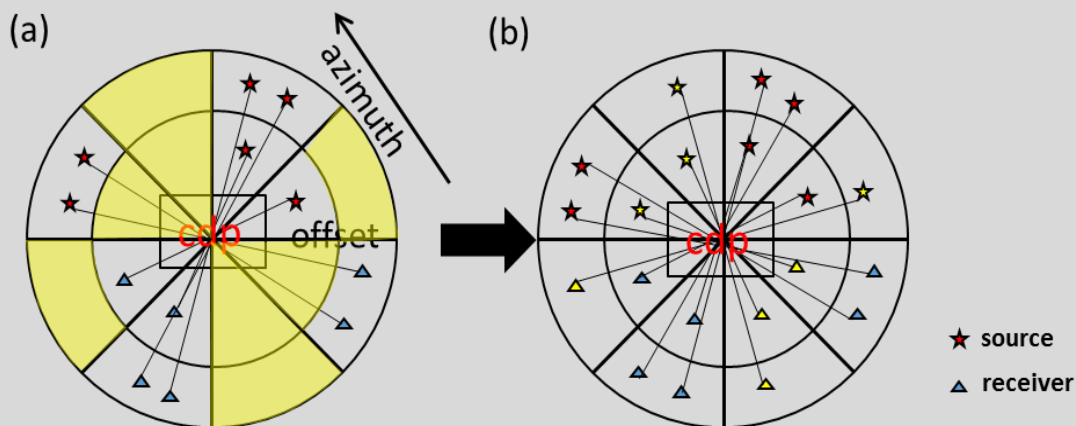
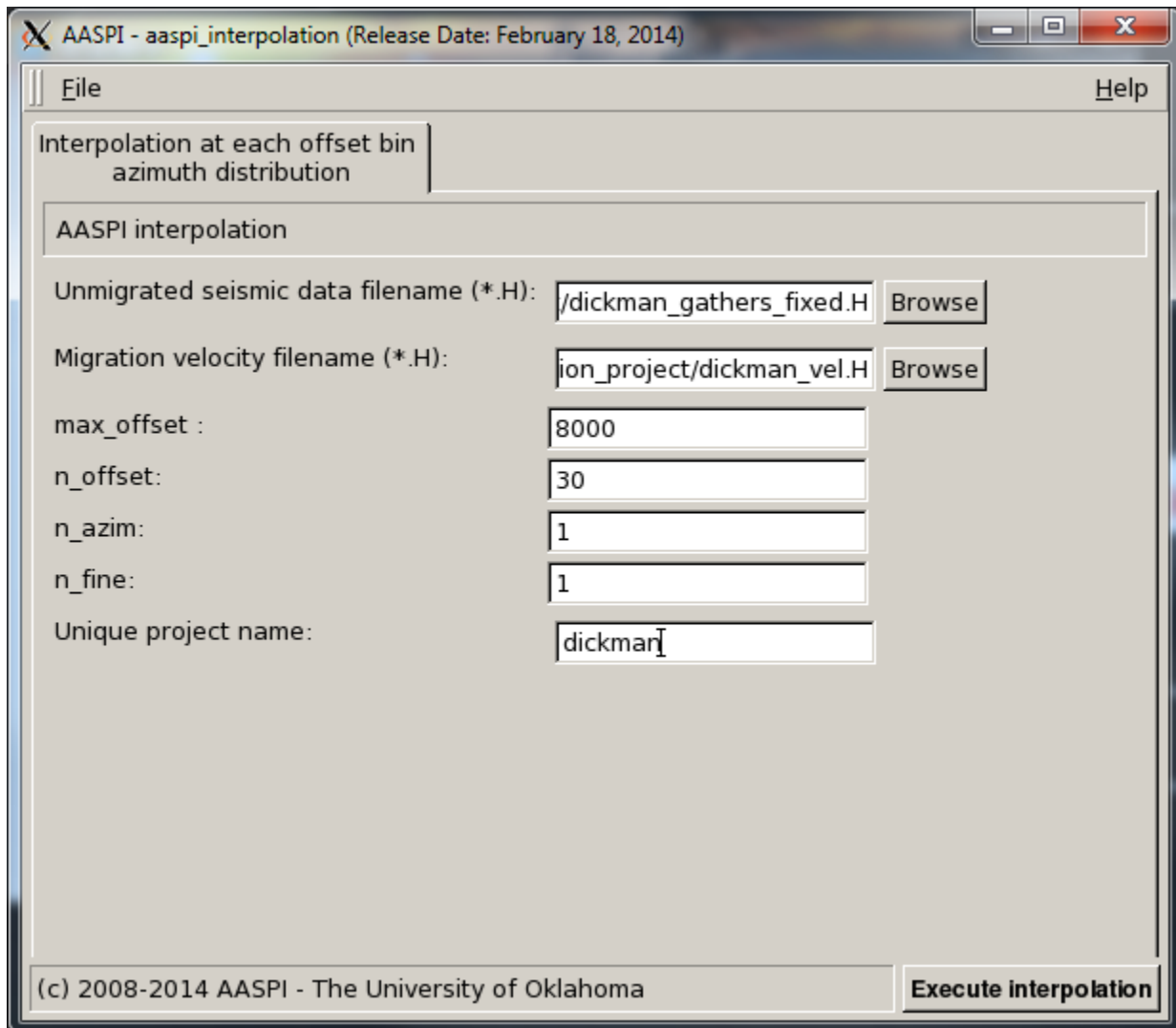


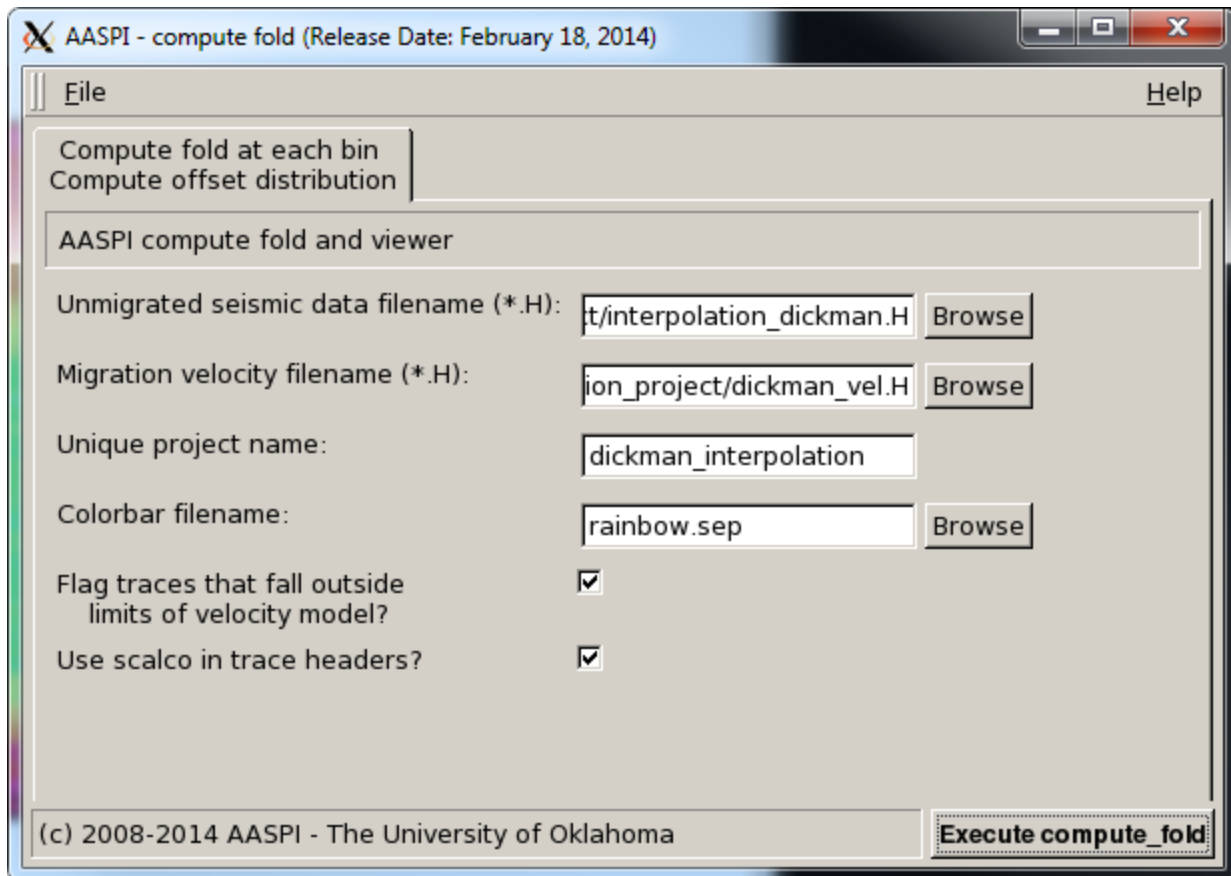
Figure 1, Cartoon showing a CMP with two offsets and four azimuthal sectors. (a) Measured data with four filled and four empty bins. (b) After interpolation each bin has at least one trace.

Computing migration-driven 5D-interpolated data

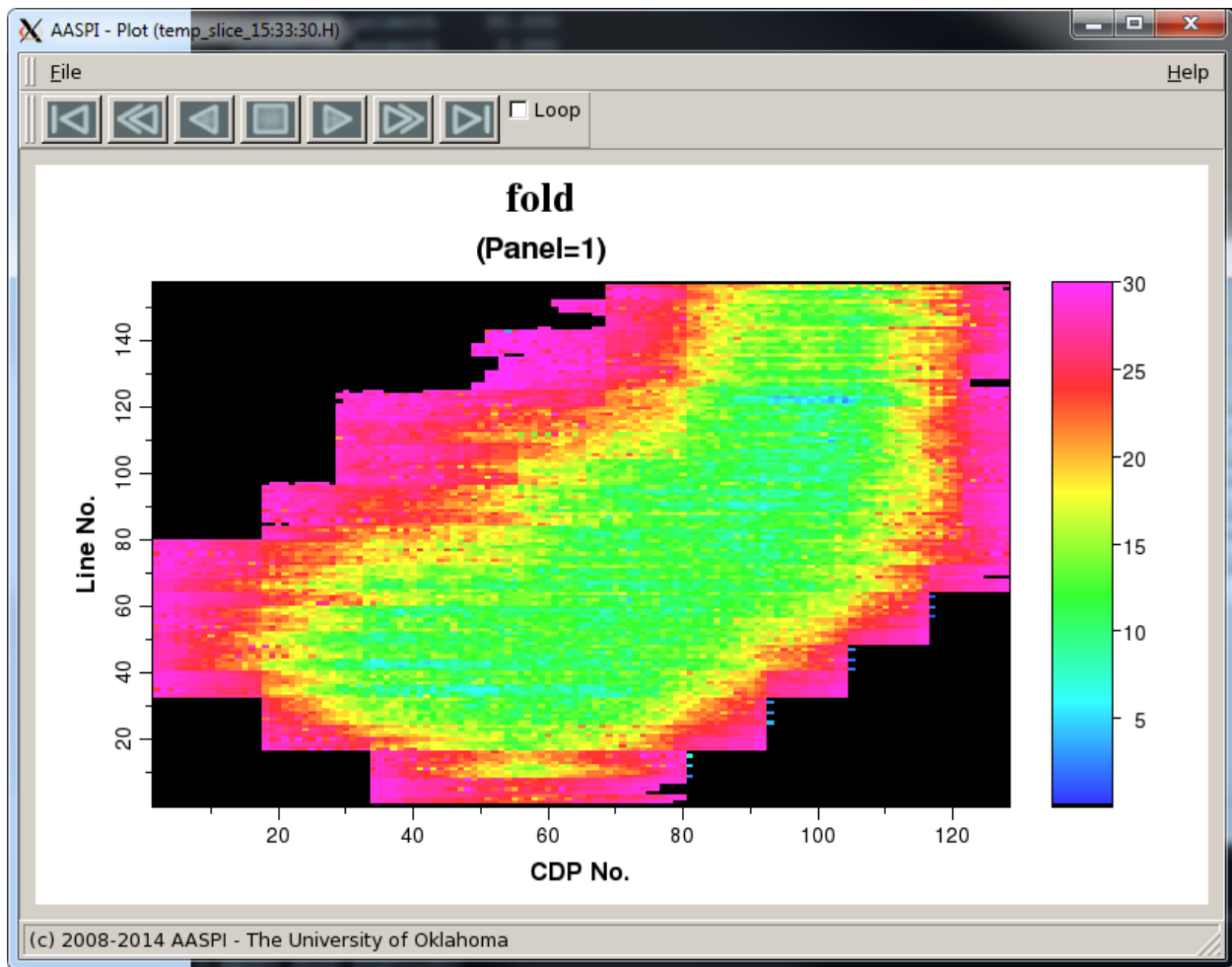


Afterward you can get file of interpolation_dickman.H, afterward please open aaspi_compute_fold as below:

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Then you can get the fold and offset distribution map of your interpolation file, then please open `aaspi_azim_offset_demig` program:

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AASPI - program azimuth_offset_demig (Release Date: February 18, 2014)

File Help

3D Kirchhoff prestack time migration into offset and azimuth bins

Input Unmigrated Seismic Gather File Name (*.H): Browse

Input Velocity Volume File Name (*.H): Browse

Input Migration File Name (*.H): Browse

Input Fold Map File Name (*.H): Browse

Input Offset Information (*.H): Browse

Mute Input (*.H): Browse

*Unique Project Name: Load default values

Suffix:

Typical Coordinates Extended

Minimum frequency:

Maximum frequency:

Number of offset bins:

Minimum Offset:

Maximum Offset:

nazim:

Taper length in time:

Max emerge angle:

Taper length for migration operator:

Migration Aperture:

t_power:

Want Migration Result? ☒ required

Output amplitude weight file? ☐

Output file showing maximum frequency imaged at each time-offset pair? ☐

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Execute azimuth_offset_demig

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Afterward you can get your interpolation file from demigration, here we name it like `mod_dickman_interpolation.H`. Then we open `aaspi_azim_offset_mig` and use `mod_dickman_interpolation.H` for migration.

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AASPI - program azimuth_offset_mig (Release Date: February 18, 2014)

File Help

3D Kirchhoff prestack time migration into offset and azimuth bins

Input Unmigrated Seismic Gather File Name (*.H): od_dickman_interpolation Browse

Input Velocity Volume File Name (*.H): ion_project/dickman_vel.H Browse

Input Fold Map File Name (*.H): d_dickman_interpolation.H Browse

Input Offset Information (*.H): t_dickman_interpolation.H Browse

Mute Input (*.H): Browse

*Unique Project Name: dickman_interpolation Load default values

Suffix: 0

Typical Coordinates Extended

Minimum frequency: 2

Maximum frequency: 120

Number of offset bins: 30

Minimum Offset: 0

Maximum Offset: 7982

nazim: 1

Taper length in time: 0.1

Max emerge angle: 30

Taper length for migration operator: 500

Migration Aperture: 10000

t_power: -1

Want Migration Result? ☒ required

Output amplitude weight file? ☐

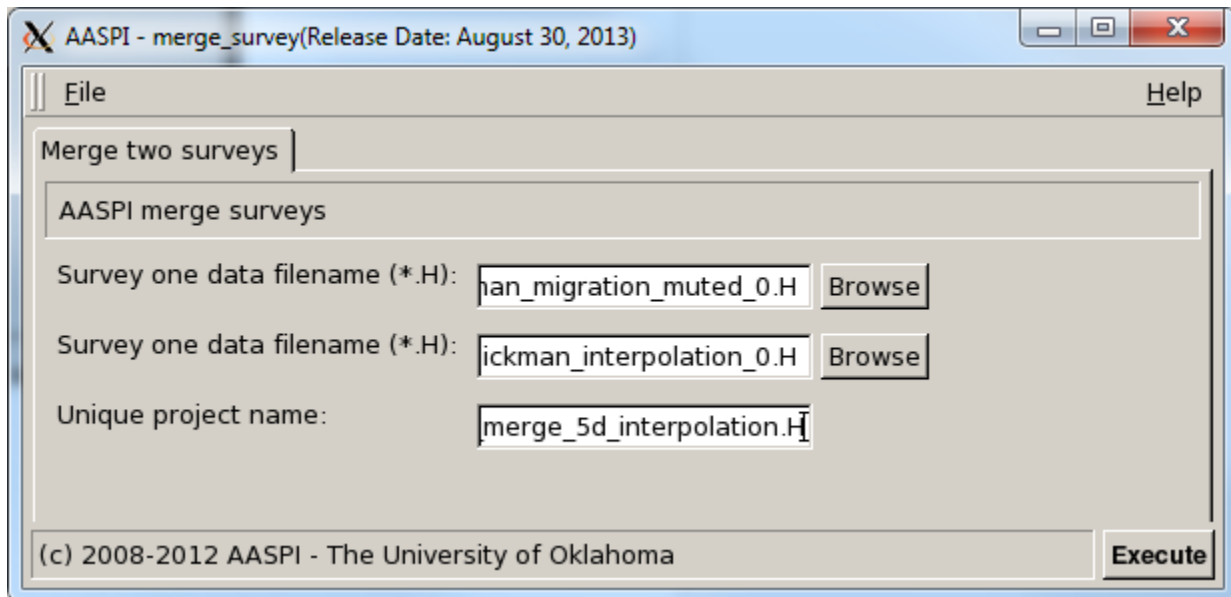
Output file showing maximum frequency imaged at each time-offset pair? ☐

Apply anti-alias filters? ☒

(c) 2008-2014 AASPI - The University of Oklahoma Execute azimuth_offset_mig

Then, you get your interpolated migrated result from your demigrated shot gathers. At last, you merge your original migrated prestack data d_mean_filt_dickman_migration_muted_0.H and interpolated migrated result d_mig_dickman_interpolation_0.H together by aaspi_merge_survey

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Ideally you can get result like this:

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